

# Localization of 3-D kinetic Alfvén wave and turbulent spectra in the solar corona region

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## ***Abstract—***

Magnetic reconnections and the Kinetic Alfvén Waves (KAW) are expected to play an important role in coronal heating but the proper methodology to describe the heating of solar corona is still being browsed by the Solar physicist. Different magnetic structures interact with the KAW to play its role in the occurring of this dramatic Corona heating. We proposed a 3D model of Kinetic Alfvén wave which is propagating in magnetic reconnection region encounters the Harris current sheet profile with taking into consideration of ponderomotive effects and non-linearity in the solar corona. This Model equation is solved numerically using the finite difference method in time and pseudospectral in spatial domain with the predictor-corrector method. The numerical simulation shows that the field structure feels a slow change without the nonlinearity whereas the presence of nonlinearity causes a rapid change. And approaching towards quasi-steady state, it generates a fully chaotic structure which are signals of turbulent filamentation with temporal evolution. We have also obtained the semi-analytical solution for these localized structures which shows the transverse scale size to be comparable to electron inertial length.

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