

Alpha Particle Confinement and Losses in JET's Tritium Campaign

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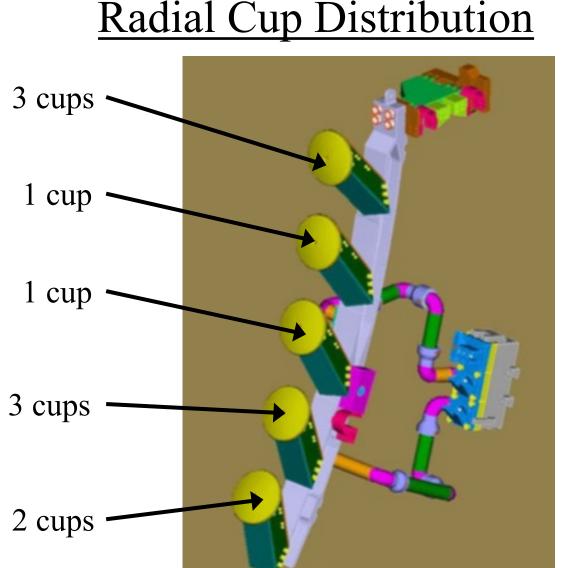
Faraday Cup Array[1]

JET's Fast Ion Loss Detectors

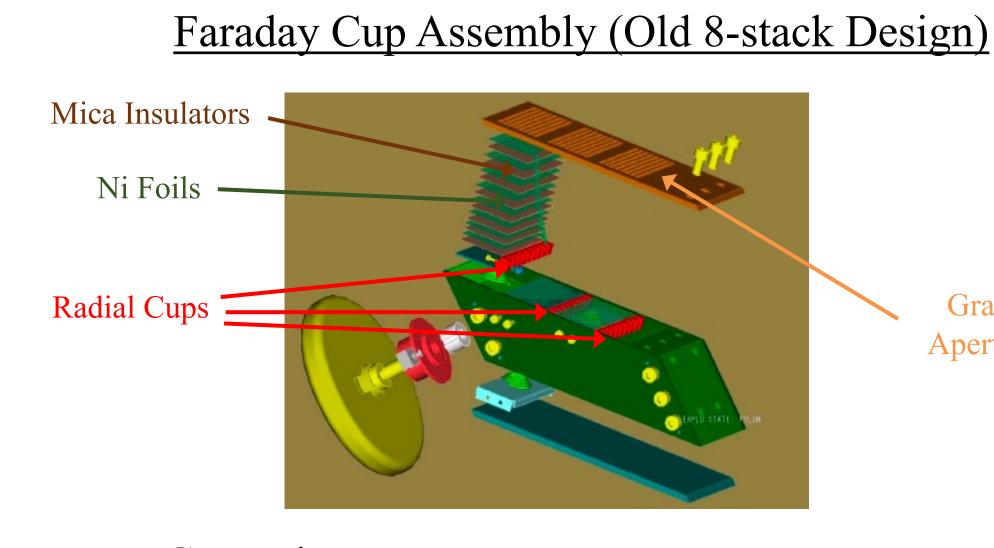
Grated

Apertures

Scintillator Probe[2]



per cup



General

- 5 pylons, up to 3 radial cups per pylon, stack of 4 foils
- Wide spatial distribution poloidally and radially near the wall
- Poloidal locations below midplane: 9°, 15°, 21°, 27°, 33°

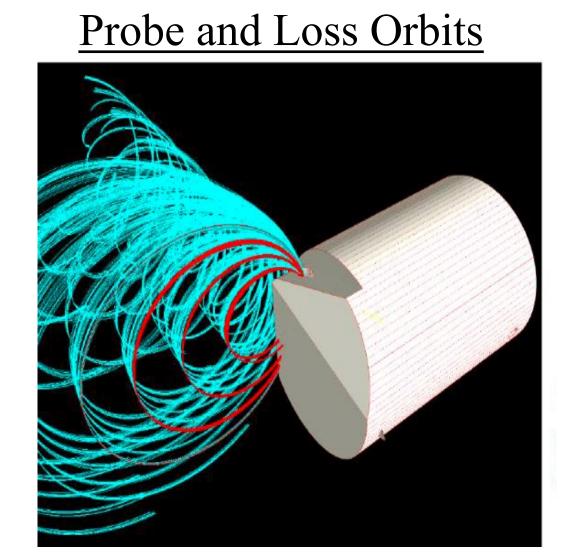
Energy Resolution per Foil^T

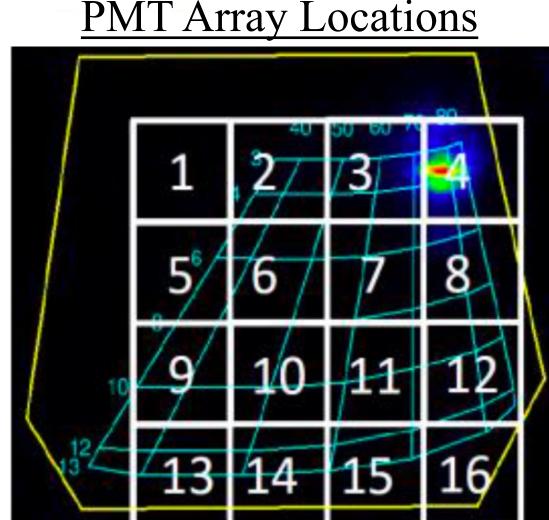
Foil	Proton (Mev)	Deuteron (Mev)	Triton (Mev)	He3 (Mev)	Alpha (Mev)
1	0.0 - 0.49	0.0 - 0.49	0.0 - 0.50	0.0 - 1.55	0.0 - 1.54
2	0.68 - 0.96	0.79 - 1.10	0.84 - 1.20	2.30 - 3.35	2.48 - 3.55
3	1.10 – 1.32	1.35 – 1.60	1.48 – 1.76	3.90 - 4.70	4.17 – 5.09
4	1.45 – 1.65	1.78 - 2.00	2.00 - 2.25	5.20 - 5.80	5.60 - 6.35

- Foil stacks are alternating layers of 2.5 μm Ni and 2.5 μm
- Ion energy and species determines deposition depth
- Measure raw current in Ni foils → Contributions from multiple species
- Nomenclature: Signal ID = Pylon #, Cup #, Foil #

^TFound via SRIM code (www.srim.org)

- e.g. $213 = 2^{nd}$ pylon from top, 1^{st} radial cup, 3^{rd} foil deep
- Rough energy resolution, pitch resolution from apertures, wide spatial resolution



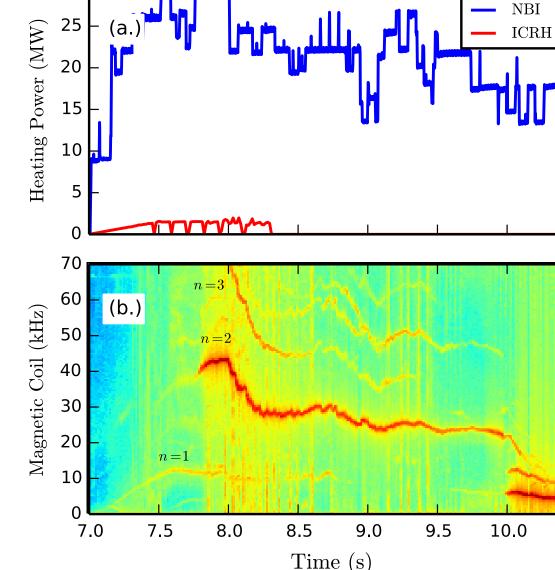


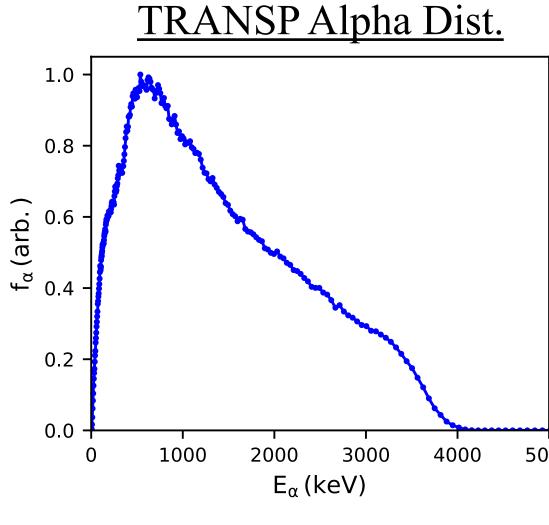
General

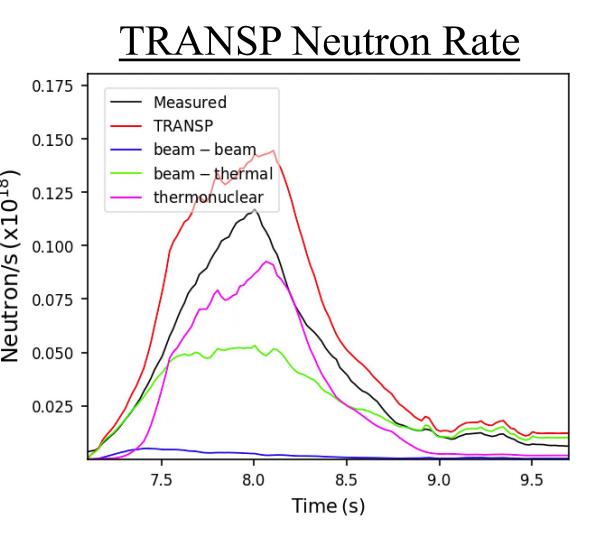
- Functions like a mass spectrometer
- Scintillator light passes through a beamsplitter on to a CCD camera and PMT array
- 16 PMTs correspond to varied zones in pitch and gyroradii (energy)
- Single location 180° toroidally opposite of Faraday cup array and poloidally consistent with the 3rd Faraday cup pylon
- Good gyroradii resolution (perp. energy), good pitch resolution, limited spatial resolution
- Predominantly sensitive to fusion products and RF-heated ions. Beam ions can appear at top edge of PMT and CCD array

Reference Discharge

External Heating and Magnetic Activity







- JET Pulse 99151: $I_p=2.3$ MA, $B_0=3.4$ T, $n_e=7\times10^{19}$ m⁻³, $T_e=8.5$ keV, 95% tritium with residual deuterium and hydrogen for minority RF heating
- Strong, low freq., MHD observed with a prolonged n=2 NTM

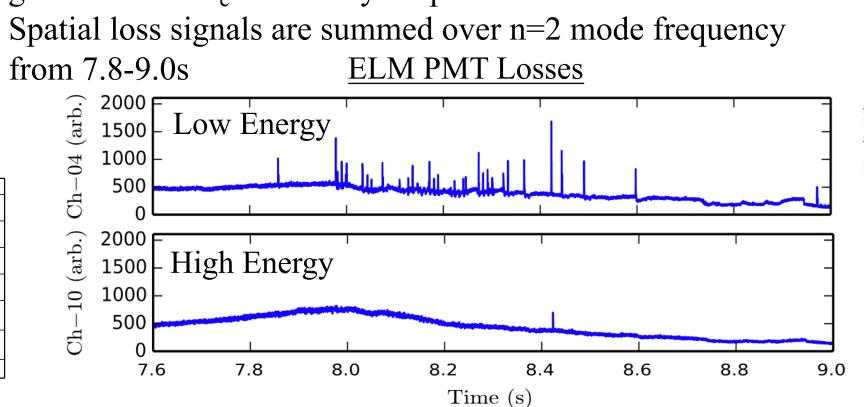
Radial

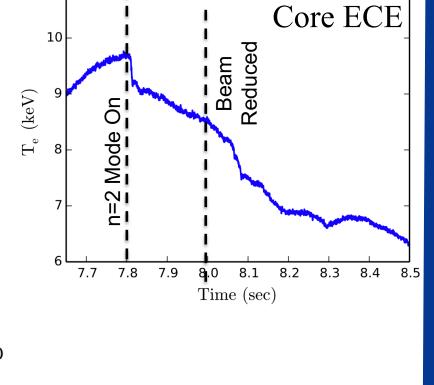
Alpha distribution from TRANSP[3] shows majority of distribution > 1 MeV and originates from thermonuclear fusion with some from beam-thermal interactions

Coherent n=2 Losses Faraday: High **Energy Losses** Probe: High **Energy Losses** 9.0 9.5 10.0 Time (s)

Measured Losses

- Faraday cup losses are corrected with the deepest foil in the stack and limit signal to the n=2 spectral feature[4]
- High energy alphas interact more with NTM while low energy alphas (and beam tritons) interact with ELM bursts
- Poloidal variation is evident and needs to be explained with modelling
- Radial losses are highest near wall and are suspected to be due to the large Larmour radii of alphas (10-14 cm). Needs computational support.
- No strong change in neutron rate due to n=2 NTM. At mode onset, neutron rate continues to grow but core T_e drastically drops

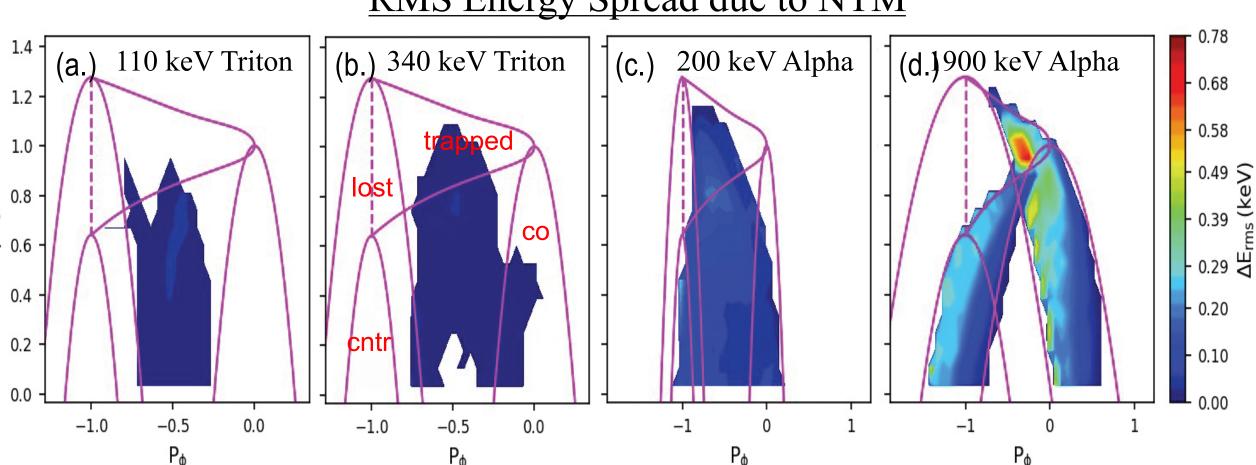




ORBIT-kick Modeling Results

- ORBIT-kick[5]: Hamiltonian guiding-center code that calculates changes in the particles' constants of motion (E - energy, P_{ϕ} - canonical momentum) due to a supplied perturbation
- An analytic form was used for NTM mode structure[6] while the mode amplitude was reasonably estimated to be $\delta b/B = 10^{-5}$
- Run Parameters: 200,00 particles for each species, 1ms runtime, electrostatic potential (from TRANSP) included, allow transitions beyond LCFS

RMS Energy Spread due to NTM

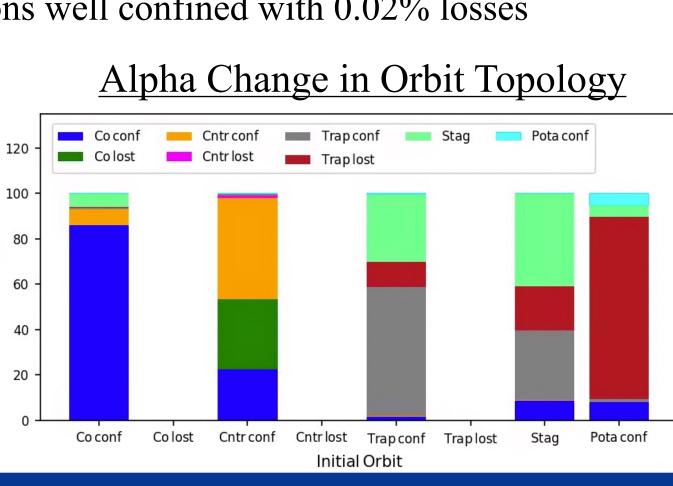


- Low energy tritons and alphas have weak interactions. Strongest interaction is for trapped alphas with energy in the MeV range
- Alphas experience 6.7% losses; tritons well confined with 0.02% losses
- Agrees with measured losses (left)

Trapped, potato, and stagnation orbits most susceptible to be lost - Region of strongest interaction 100 above

measurement

- FILDs most likely to detect lost trapped orbits, so in good agreement with



Ongoing and Future Work

- Compare measured spatial loss distribution to that obtained with ORBIT
- Further constrain with synthetic diagnostic simulations[7]
- Investigate any alpha effects related to heating and neutron production
- Interpret and compare to alpha losses in the DT-campaign
- Extend DT scenarios to ITER to estimate alpha transport and losses due to low freq. MHD



Poloidal

References

- ¹P. J. Bonofiglo, V. Kiptily, A. Horton, P. Beaumont, R. Ellis, F. E. Cecil, M. Podesta, and JET Contributors 2020 Rev. Sci. Instrum. 91 093502 ²S. Baeumel, A. Werner, R. Semier, S. Mukherjee, D. S. Darrow, R. Ellis, F. E. Cecil, L. Pedrick, et al. 2004 Rev. Sci. Instrum. 75 3563 ³B. Joshua, G. Marina, P. Francesca, S. Jai, P. Alexei, and P. Gopan 2018 TRANSP Software USDOE Office of Science (SC), Fusion Energy Sciences (FES) (https://doi:10.11578/dc.20180627.4)
- ⁴P. J. Bonofiglo et al. "Lost Alpha Faraday Cup Foil Noise Characterization During JET Plasma Post-Processing Analysis" Rev. Sci. Instrum. (Under Review)
- ⁶M. Gobbin et al. 2009 Nucl. Fusion **49** 095021 ⁷P. J. Bonofiglo et al. 2022 *Nucl. Fusion* **62** 026026

⁵M. Podesta, et al. 2017 Plasma Phys. Control. Fusion **59** 095008

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